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The new JENSA gas-jet target for astrophysical radioactive beam experiments

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Abstract

To take full advantage of advanced exotic beam facilities, target technology must also be advanced. Particularly important to the study of astrophysical reaction rates is the creation of localized and dense targets of hydrogen and helium. The Jet Experiments in Nuclear Structure and Astrophysics (JENSA) gas-jet target has been constructed for this purpose. JENSA was constructed at Oak Ridge National Laboratory (ORNL) where it was tested and characterized, and has now moved to the ReA3 reaccelerated beam hall at the National Superconducting Cyclotron Laboratory (NSCL) at Michigan State University for use with radioactive beams.

1 The application of nuclear transfer reaction studies to exotic nuclei is
2 a primary method by which reactions of astrophysical interest can be ex-
3 perimentally constrained. Such studies require that the outgoing reaction
4 products be detected with good energy and angular resolution and that the

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5 reactions of interest be clearly distinguished from background reactions. Ap-
6 plication of transfer reactions techniques on exotic nuclei are greatly facili-
7 tated by the availability of a localized high-density target of gaseous elements.
8 The target must be of sufficient density ($> 5 \times 10^{18}$ atoms/cm²) to apply these
9 techniques to exotic beams. For a given elemental thickness, the use of pure
10 gases reduces the energy loss of the beam through the target (i.e., degra-
11 dation of the energy resolution is minimized) and reduces the probability
12 of contaminant reactions. Ideally, the gas volume should be well-matched to
13 the beam spot size (~ 3 mm) such that reaction angles can be unambiguously
14 determined. Finally, the beam-target interaction point should be surrounded
15 with an efficient high-resolution charged particle array [e.g., the SuperOR-
16 RUBA (Oak Ridge Rutgers University Barrel Array)[1] of silicon detectors].
17 The JENSA gas-jet target has been constructed with these design criteria
18 in mind [2] allowing approximately 50% coverage over the detector angular
19 range $20^\circ - 160^\circ$ in the laboratory.

20 The operating principle of the JENSA target is as follows. Target gas is
21 compressed by an industrial compressor [PDC Machines PDC-4-100-500(150)]
22 to high pressure (200-400 psi) before injection through a Laval nozzle into
23 the target chamber. After traveling ~ 14 mm through vacuum, the gas is
24 collected in a pair of concentric conical receivers which are pumped in par-
25 allel by high throughput roots pumps (Leybold WSU2001 and WSU1001).
26 Multistage rootsblowers (Ebara A10S) compress the gas further before injec-
27 tion back into the PDC compressor for recirculation. Differential pumping
28 through flow-restricting apertures has been employed to maintain low beam-
29 line pressures outside of the target. The aperture sizes range from 3-5 mm
30 and 15-28 mm on the upstream and downstream sides of the target chamber,
31 respectively. Each pumping region, separated by an aperture, is indepen-
32 dently pumped by turbomolecular pumps resulting in pressures ranging from
33 $\sim 10^{-6} - 10^{-7}$ Torr at 50 cm from the target jet. The beam is tuned through
34 these apertures for bombardment of the target gas.

35 The spatial distribution and density of the gas jet have been determined
36 with measurements of α energy loss from radioactive sources [2]. One detec-
37 tor of the SuperORRUBA array was used to detect the α particles after they
38 traveled through the gas jet. The transmission of the α particles through
39 the jet produces energy loss observable from the energy vs. position detector
40 spectra. The energy loss measured in the shadow of the jet could be ex-
41 trapolated back to the jet position to determine the target width and areal
42 density. For a jet of ⁴He gas injected through a Laval nozzle 1.1-mm wide

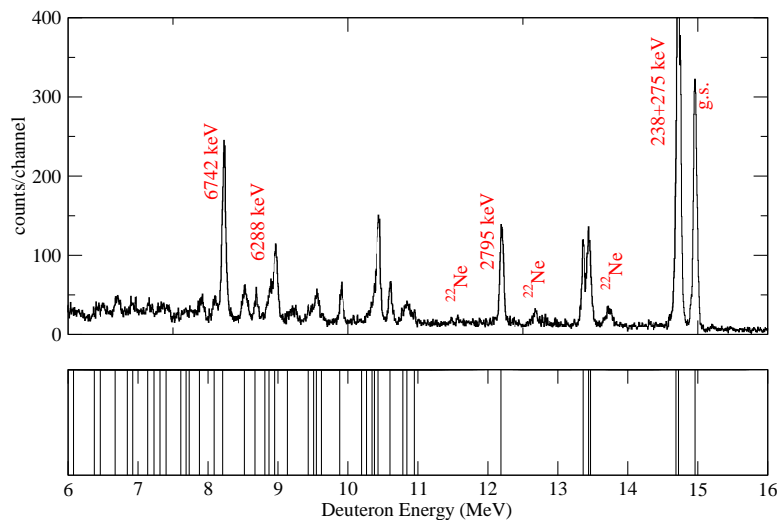


Figure 1: (top) The deuteron spectrum from the $^{20}\text{Ne}(p, d)^{19}\text{Ne}$ reaction are shown. All observed peaks are from reactions on ^{20}Ne except for those labeled ^{22}Ne . (bottom) The expected energies of deuterons populating known levels in ^{19}Ne .

43 at an inlet pressure of 400 psi, target densities greater than 10^{19} atoms/cm²
 44 were obtained with a spatial width of 5 mm FWHM.

45 First experiments with JENSA have been performed at the ORNL Ho-
 46 lifield Radioactive Ion Beam Facility (HRIBF). Targets were made of natural
 47 abundance He, N₂, and Ne gases. These were used to study the $^4\text{He}(^{15}\text{N}, \alpha)^{15}\text{N}$,
 48 $^{14}\text{N}(p, t)^{12}\text{N}$, $^{20}\text{Ne}(p, d)^{19}\text{Ne}$, and $^{20}\text{Ne}(p, t)^{18}\text{Ne}$ reactions to address open
 49 questions in nuclear astrophysics and nuclear structure. The observed deuteron
 50 energy spectrum from the $^{20}\text{Ne}(p, d)^{19}\text{Ne}$ study is plotted in Fig. 1. Excellent
 51 correspondence was obtained between the observed spectrum and known en-
 52 ergy levels in ^{19}Ne indicating the chemically-pure nature of the target. This
 53 is in sharp contrast to the previous attempt to study this reaction with an
 54 implanted target, which suffered from severe contamination of the spectrum
 55 due to the large number of the target constituents [3, 4].

56 In the Fall of 2013, the JENSA target was moved to the reaccelerated
 57 beam hall, ReA3 [5], at the NSCL for use with exotic beams. At ReA3
 58 primary beam fragments are thermalized in a gas cell, extracted, and then
 59 reaccelerated to energies of 3-6 MeV/u. The JENSA target will be used to
 60 study reactions in inverse kinematics such as $(^3\text{He}, d)$, (d, p) , and (α, p) on
 61 beams primarily of astrophysical interest. Commissioning of JENSA at ReA3
 62 is in progress but stable beams have already been successfully tuned through

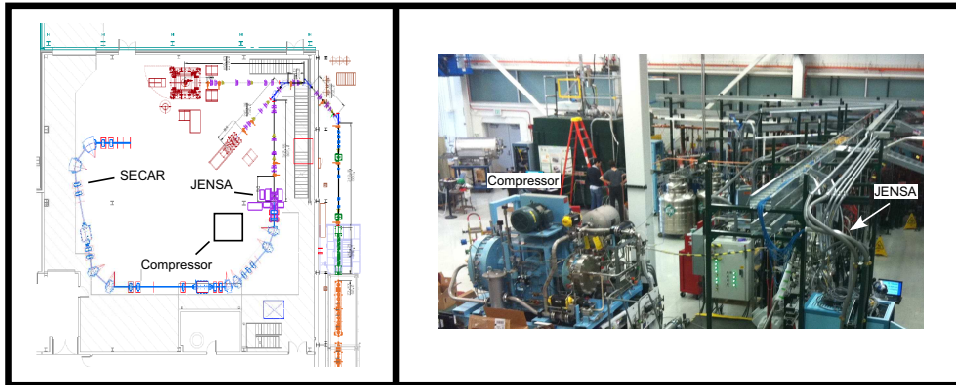


Figure 2: (right) A picture showing the placement of JENSA in the ReA3 hall. The industrial compressor is visible at the bottom left. (left) A schematic of the ReA3 experimental hall. A preliminary design of the coupling to SECAR is shown.

63 the target. Use of the target for the study of $({}^3\text{He}, d)$ reactions is of high
 64 priority, and approximately one-half of the required 200 liters of ${}^3\text{He}$ has been
 65 procured with the remainder expected to be purchased in 2016. A picture of
 66 JENSA in place at ReA3 is shown in Fig. 2. As the NSCL transitions to the
 67 Facility for Rare Isotope Beams, it is envisioned that JENSA will become
 68 the target for a new recoil separator SECAR [6].

69 In conclusion, advances in nuclear astrophysics are often driven by tech-
 70 nical advances in studying the nuclear reaction rates. The construction of
 71 the JENSA gas-jet target represents such an advance and enables a new pro-
 72 gram of studies with exotic beams. First tests and experiments with JENSA
 73 at HRIBF have demonstrated its remarkable performance. JENSA has now
 74 been installed at ReA3 at the NSCL and commissioning is in progress. The
 75 first measurement with a radioactive beam will be a study of the ${}^4\text{He}({}^{34}\text{Ar}, p){}^{37}\text{K}$
 76 reaction of importance to X-ray burst nucleosynthesis [7]. This work was sup-
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